

Towards an Integrated Geomagnetic Polarity Reversal Timescale for the Pleistocene

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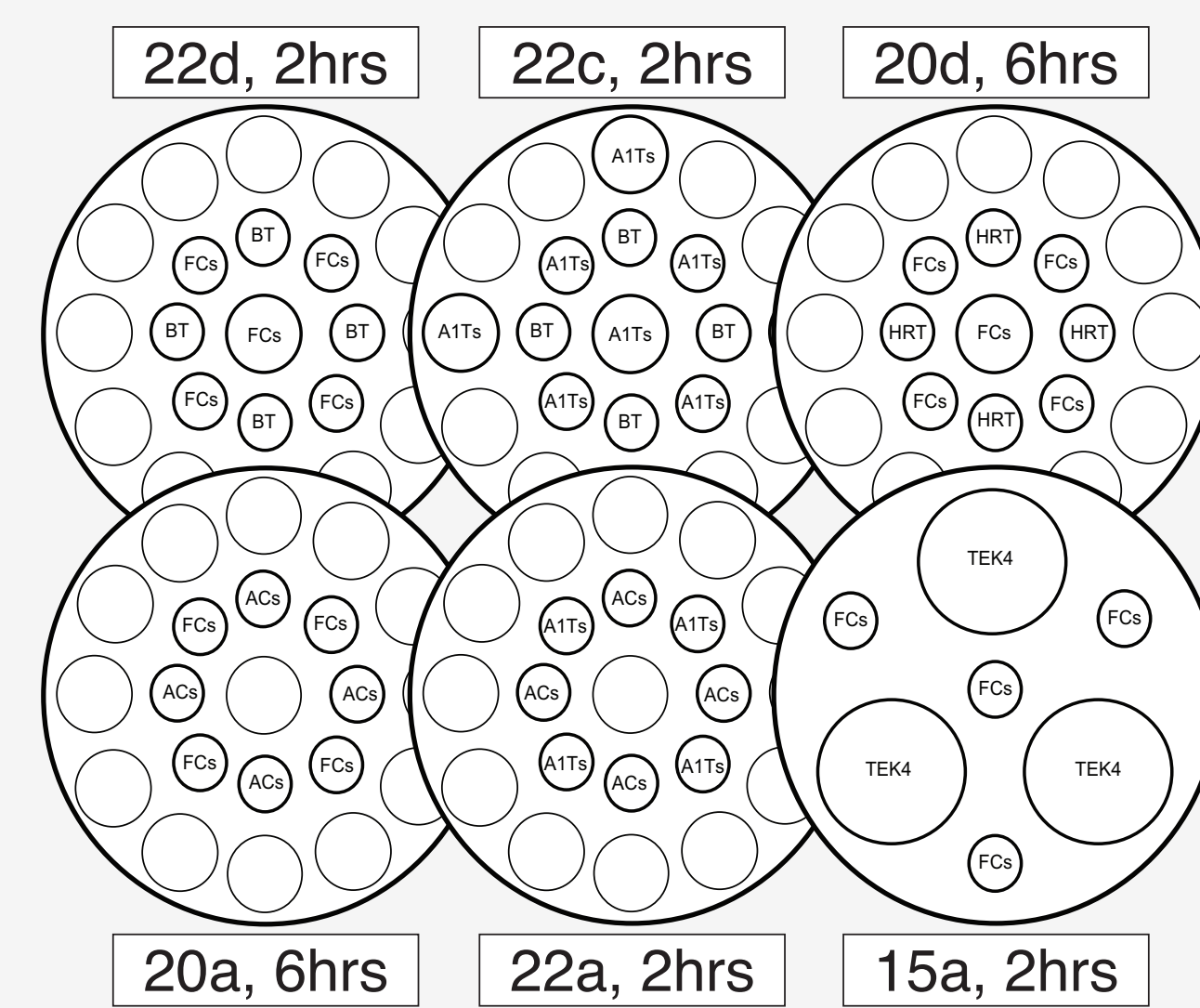


Abstract

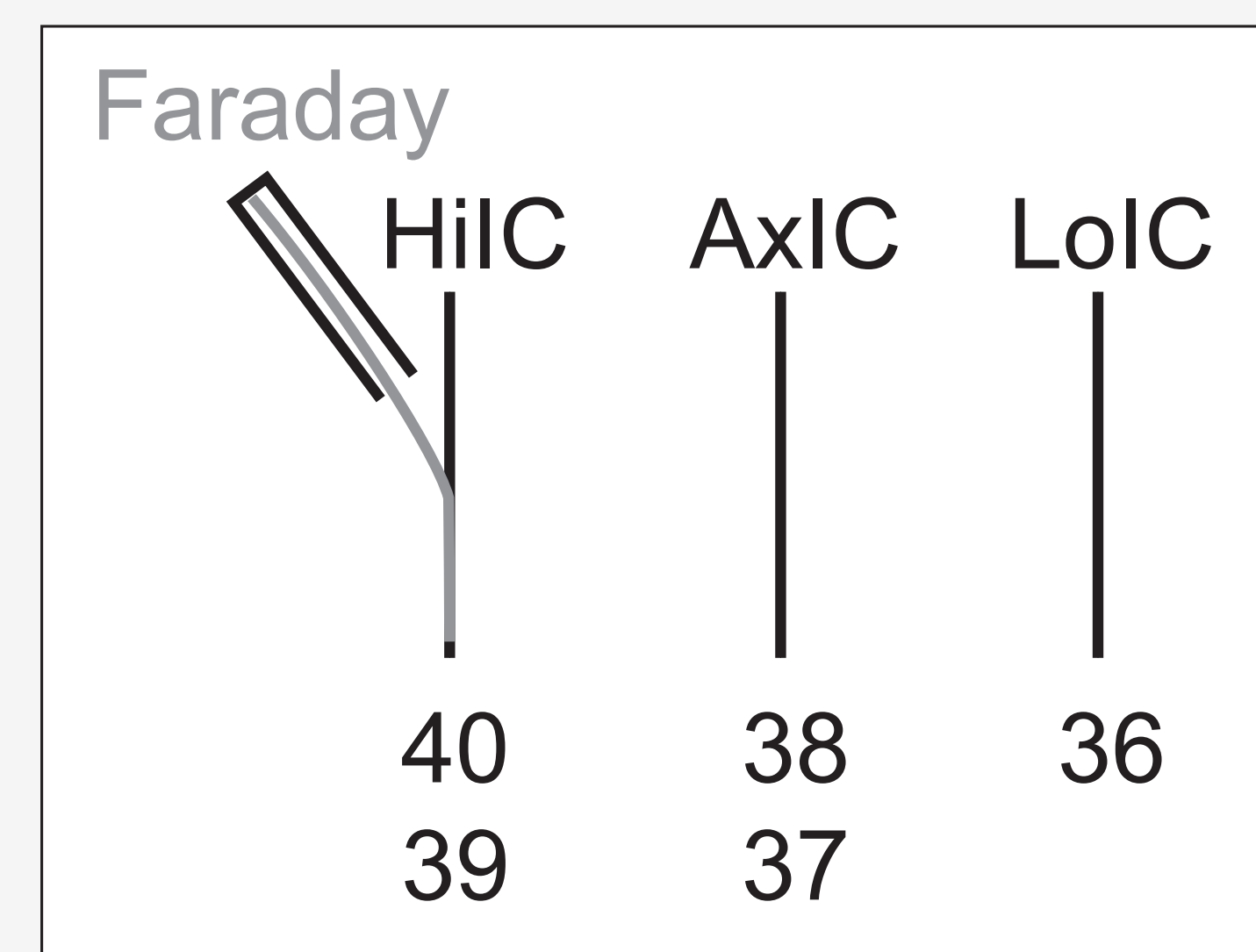
The development of the geomagnetic polarity timescale (GPTS) in the mid 20th century led to the greater understanding of seafloor spreading and plate tectonics (Heirtzler et al., 1968). Over 40 years later, the GPTS continues to be refined, particularly in terms of integrating multiple dating techniques to improve precision of such events, or to resolve the duration of geomagnetic transitions. Recent advancements in integrating astronomical and ⁴⁰Ar/³⁹Ar dating techniques, and improving upon the precision of neutron fluence monitors, necessitate re-evaluation of the accuracy and precision of various geologic events. Here, we review the ages of three Pleistocene geomagnetic polarity reversals: the Matuyama-Brunhes (ca. 0.78 Ma), the Cobb Mountain (ca. 1.2 Ma), and the Reunion (ca. 2.1 Ma) events. High-precision astronomically calibrated ⁴⁰Ar/³⁹Ar ages have been obtained via a Noblesse multi-collector noble gas mass spectrometer on volcanic and other datable materials related to each event. The ages were derived by single- or multi-crystal total fusion and/or step heating experiments, using the astronomically calibrated Fish Canyon sanidine and/or the astronomically tuned A1 sanidine as monitor minerals. Each of these ages is then compared to independent astronomical ages for the events in order to define tie-points for constructing a Pleistocene a multi-chronometer GPTS. Although only three reversals are addressed here, the methodology applied shows promise to refining short-lived excursions to enable further understanding of the wavering magnetic field.

Methods: ⁴⁰Ar/³⁹Ar technique

Irradiation

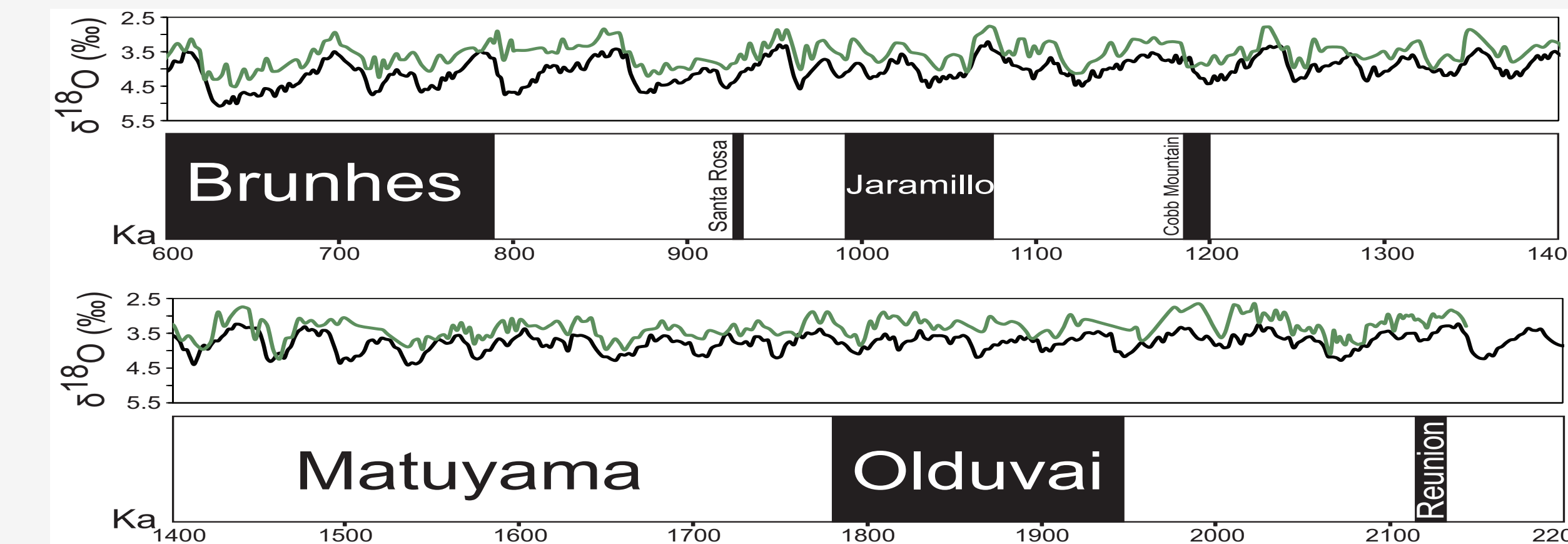


Noblesse Configuration

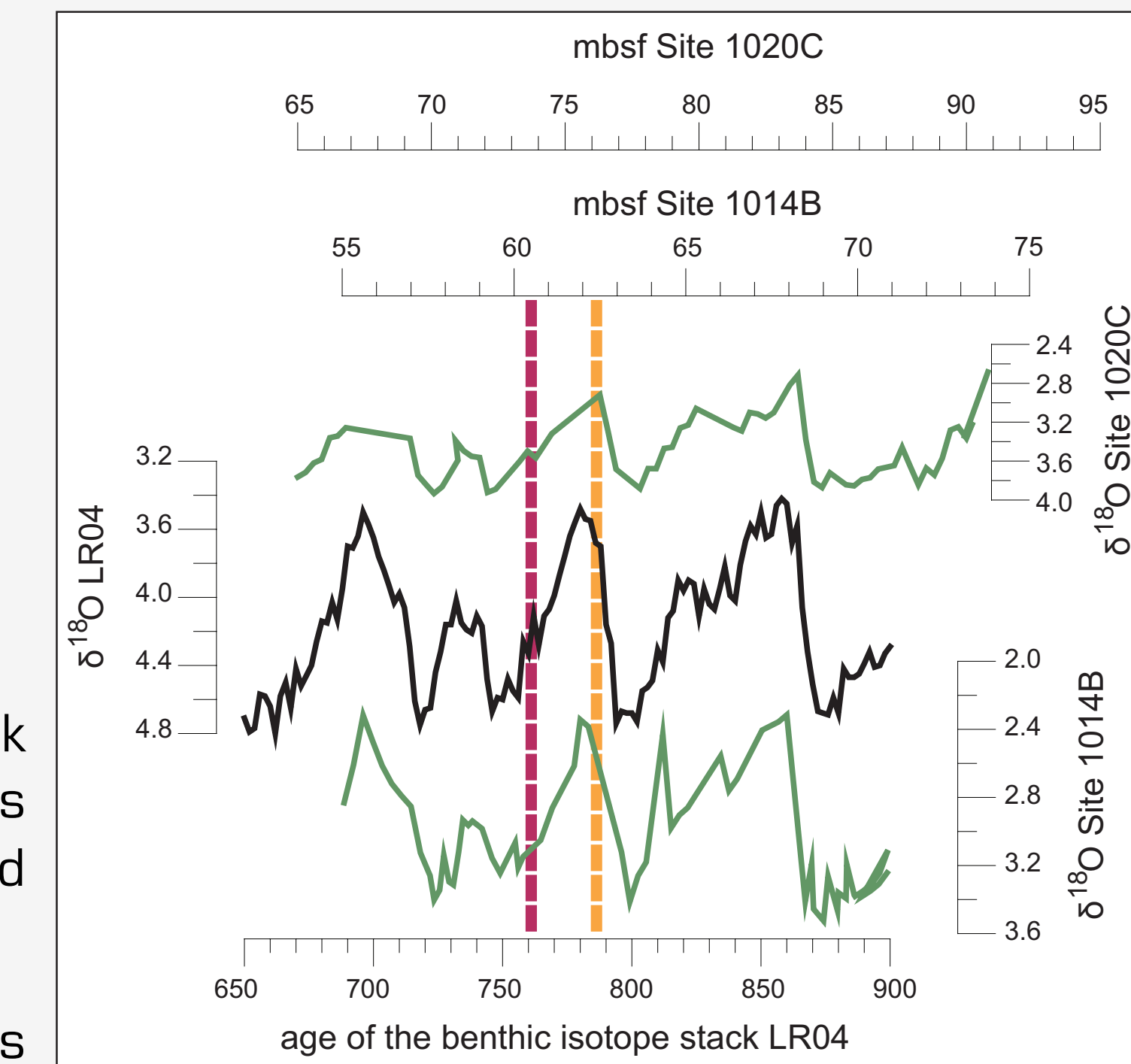


Methods: astronomical tunings

MD972143, Philippine Sea (after Horg et al., 2002)



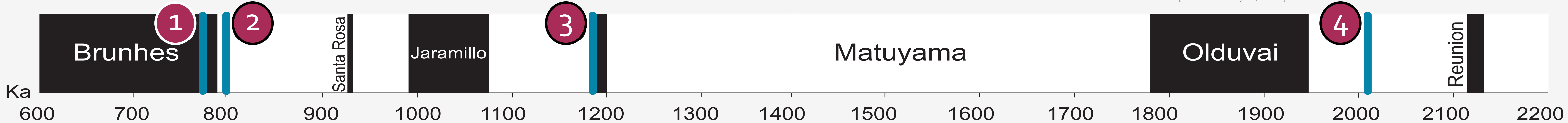
ODP Leg 167, California Margin



$\delta^{18}O$ record from drill cores (green) matched to the Lisiecki and Raymo global stack (black), coupled to magnetostratigraphy (ODP record from core 1020C). The positions of the Bishop Tuff (red) and M-B reversal (orange) in the ODP cores were determined by depth in core. **Bishop Tuff confirmed through EMP analyses**

All astronomically-derived ages incur a minimum of ± 1 ka uncertainty for inaccuracies in the tuning; ages < 1 Ma incur an additional ± 4 ka and ages between 1 and 3 Ma incur an additional ± 6 ka for uncertainties in LRO4 (Lisiecki and Raymo, 2005)

Setting Multi-isotopic Anchors for the Pleistocene GPTS



1 Bishop Tuff, California
Normal polarity, Brunhes chron
Matuyama-Brunhes reversal

2 Australasian Tektite, Thailand
Reverse polarity, Matuyama chron
Matuyama-Brunhes reversal

3 Alder Creek Rhyolite, California
Normal polarity, Matuyama chron
Cobb Mountain excursion

4 Huckleberry Ridge Tuff, Idaho
Normal polarity, Matuyama chron
Reunion excursion

Matuyama-Brunhes reversal in the Pacific

- Astronomical age of 0.787 ± 0.005 Ma (MD core)
 - Astronomical age of $0.785 - 0.789 \pm 0.008$ Ma (ODP core)
 - Astronomical age of 0.773 for Atlantic sites (Channell et al., 2010)
 - ΔT between reversal and Bishop Tuff averages 0.020 ± 0.003 Ma (based on sedimentation rates determined in this study)
 - ΔT between reversal and Tektite is 0.0110 ± 0.0085 Ma (this study)
- Integrated (⁴⁰Ar/³⁹Ar, U-Pb) Age of 0.787 ± 0.004 Ma**

Cobb Mountain excursion

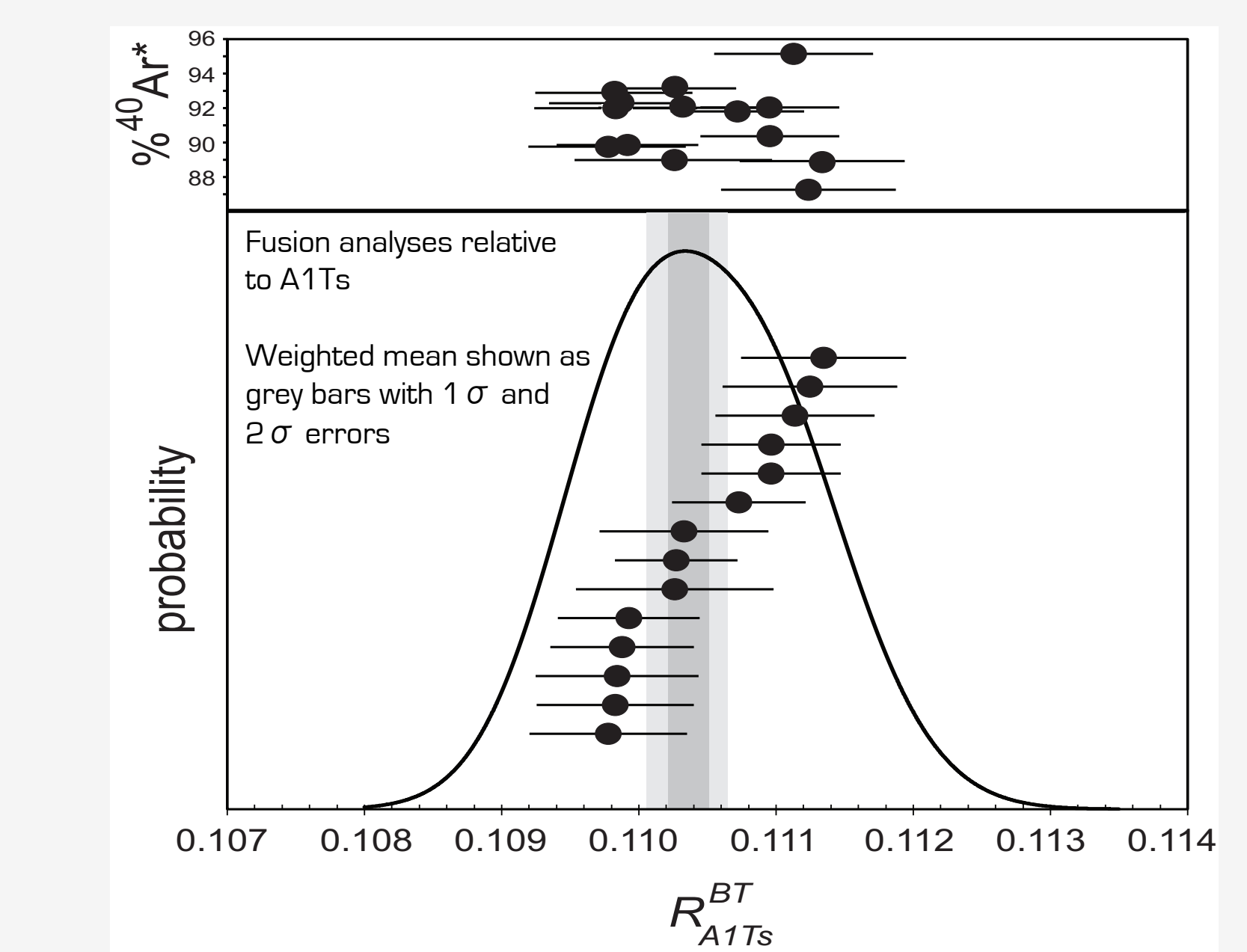
⁴⁰Ar/³⁹Ar Age of 1.184 ± 0.001 Ma

Reunion excursion(s)

- Reunion I astronomical age of $2.115-2.131$ Ma (MD core)
- **Integrated (⁴⁰Ar/³⁹Ar, K-Ar) Age of 2.0767 ± 0.0020 Ma**
- Supports notion of 2 distinguishable Reunion events?
- Or dating younger Member C with antecrystic inheritance? (see Ellis et al., poster V53B-2610, Friday afternoon)

References and Acknowledgements

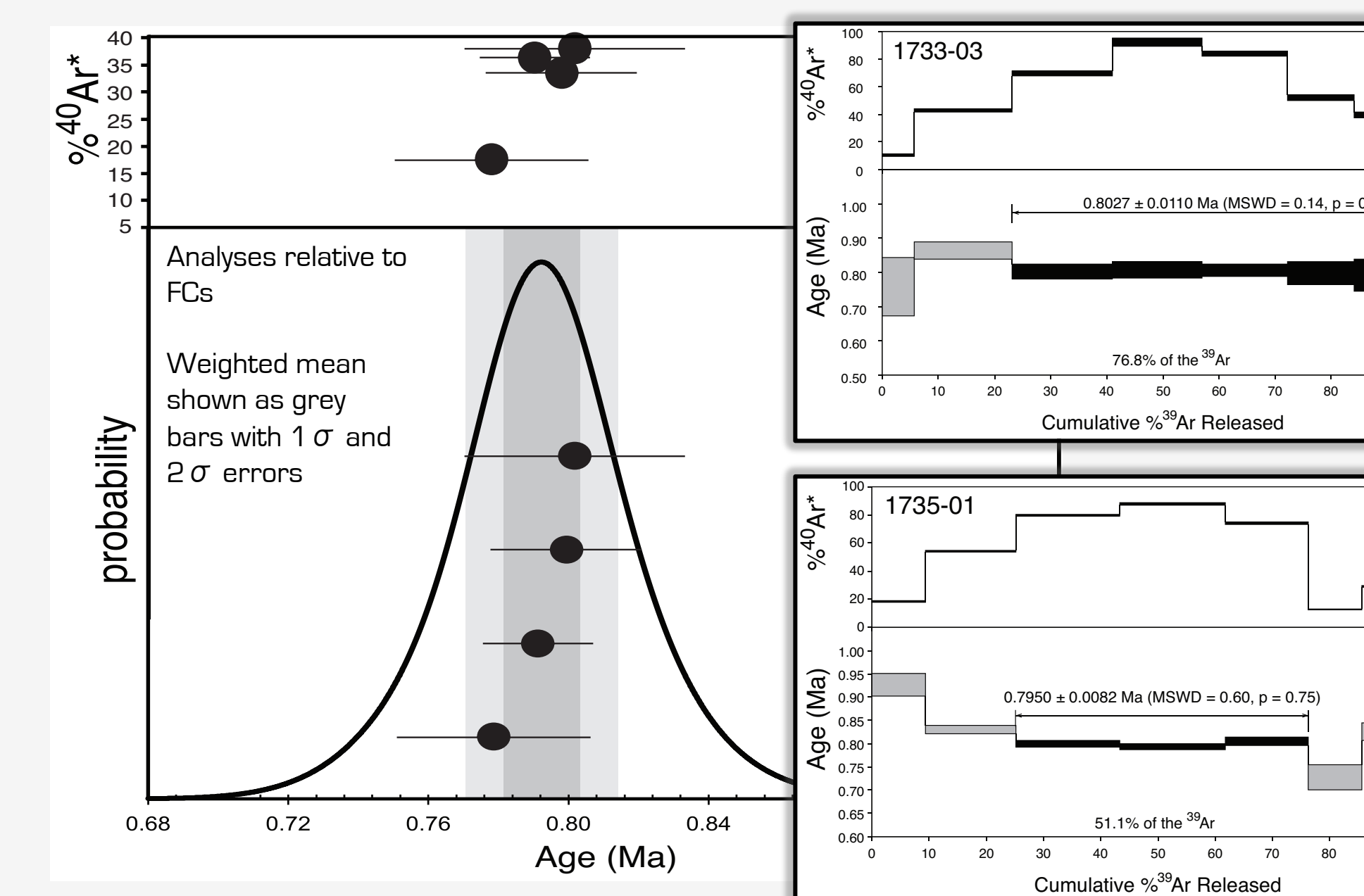
Channell et al., 2010 G-Cubed; Crowley et al., Geology 2007; deMenocal & Baker, Proc ODP Sci Res 2000; Heirtzler et al., JGR 1968; Horg et al., EPS 2002; Lisiecki & Raymo, Paleocool. 2005; Quidelleur et al., Geophys. J. Int. 2010; Rivera et al., EPSL 2011. The research leading to these results has received funding from the European Community's 7th Framework Programme [FP7/2007-2013] under grant agreement no. 215458. The QuadLab is funded by the Villum Foundation.



⁴⁰Ar/³⁹Ar Age of 0.7674 ± 0.0022 Ma
0.29%, 2σ, MSWD = 1.09, n = 14
A1T1s = 6.943 ± 0.005 Ma, Rivera et al., 2011

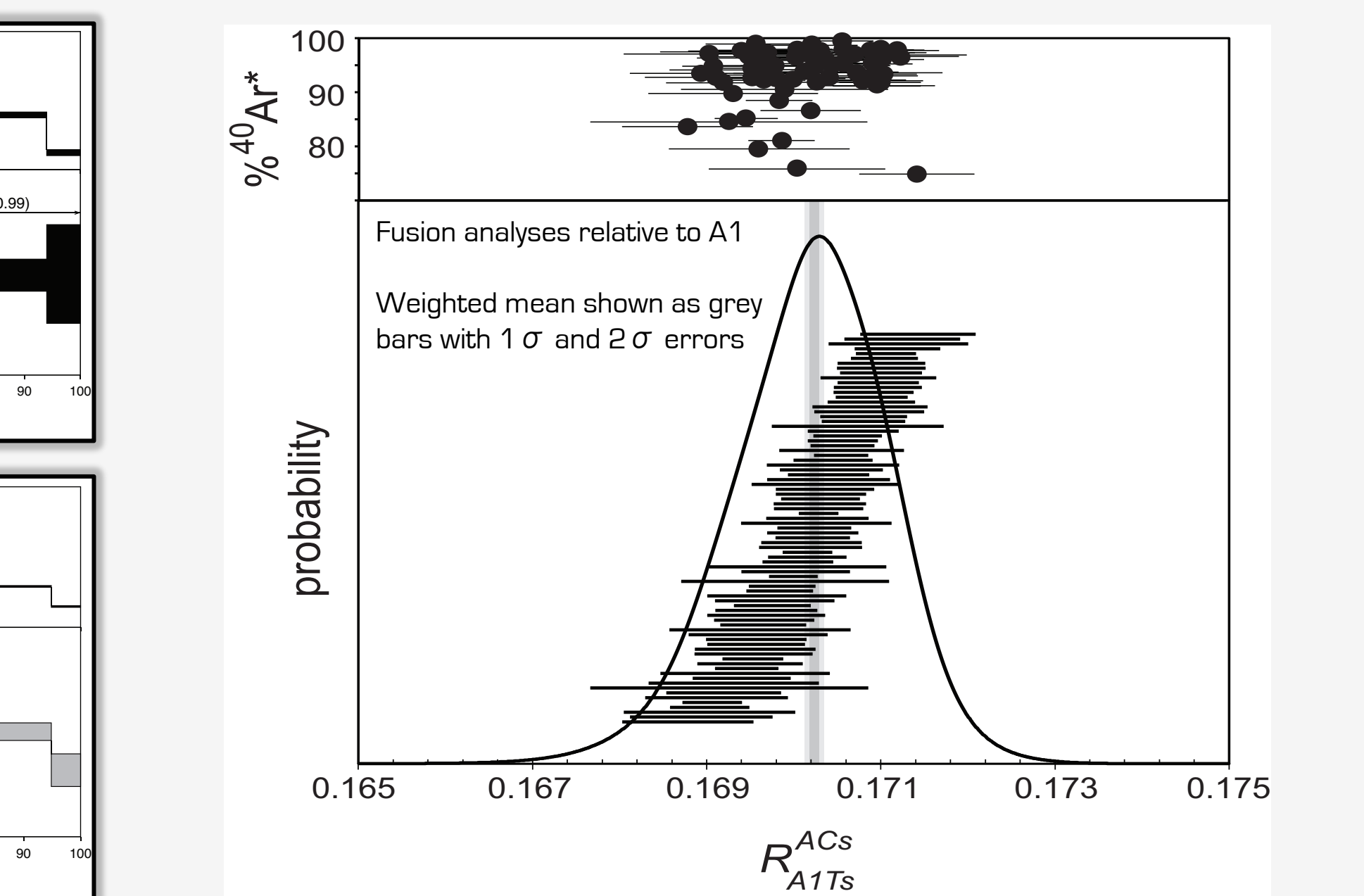
Astronomical Age of 0.765 ± 0.005 Ma
Derived from ODP core tuning

U-Pb Age of 0.7671 ± 0.0019 Ma
Crowley et al., 2007



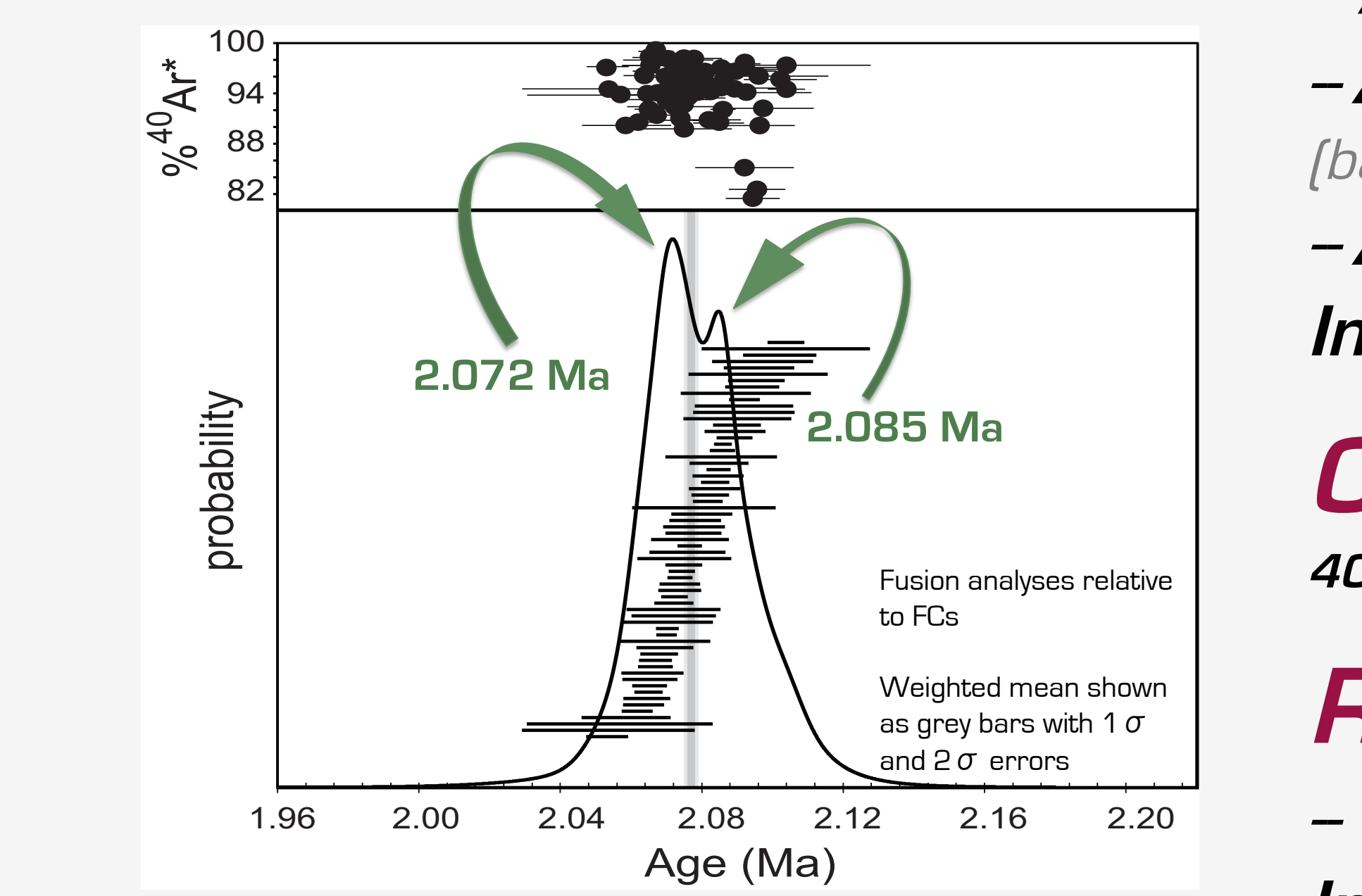
⁴⁰Ar/³⁹Ar Age of 0.7973 ± 0.0062 Ma
0.32%, 2σ, MSWD = 0.32, n = 13/20
FCs = 28.172 ± 0.018 Ma, Rivera et al., 2011

Astronomical Age of 0.798 ± 0.005 Ma
Derived from MD core tuning



⁴⁰Ar/³⁹Ar Age of 1.1839 ± 0.0012 Ma
0.10%, 2σ, MSWD = 1.38, n = 81/89
A1T1s = 6.943 ± 0.005 Ma, Rivera et al., 2011

Astronomical Age of 1.187 ± 0.007 Ma to 1.199 ± 0.007 Ma
Derived from MD core tuning



⁴⁰Ar/³⁹Ar Age of 2.0771 ± 0.0020 Ma
0.10%, 2σ, MSWD = 3.17, n = 63/73
FCs = 28.172 ± 0.018 Ma, Rivera et al., 2011

K-Ar Age of 2.04 ± 0.02 Ma
Derived from groundmass analyses of Reunion basalt samples, Quidelleur et al., 2010